Photomultiplier (PMT) basic elements



PMT formats



PMT formats range from 1/2"dia. tubes to 60 cm and more, window shape may be circular, hemispherical or hexagonal number of dynodes is from 6 to 12 (max 14) with any S-type photocathode to cover the 150 to 1200 nm spectral range

Typical values

Typically, at 200 V of inter-dynode voltage, the dynode gain

is g=4 per stage, the transit time is 2.5 ns per stage and its

time-spread is 0.5 ns

For example, with 12-stage PMTwe have:

- a *gain*, $G = 4^{12} = 1.6 10^{7}$
- an (unessential) pulse *delay*, 2.5x12 = 30 ns
- a pulse *duration* due to spread, $\tau = 0.5 \text{ x} \sqrt{12} = 1.8 \text{ ns}$ (the impulse response)
- a *current peak*, $I = eG/\tau = 1.6 \ 10^{-19} \ 1.6 \ 10^7 / \ 1.8 \ 10^{-9} = 1.4 \ mA$ (well detectable)

Single Electron Response (SER)



Secondary multiplication



For one primary electron impinging on a dynode with energy E_0 we find n secondary electrons emitted.

Number is not fixed, varies from primary to primary.

Probability of having n secondaries is Poisson distributed, $p(n) = g^n e^{-g} / n!$ where g is the mean value (or mean gain of the dynode)

Dynode Materials



The electron optics



Excursus: Design of the electron optics

Given a tentative structure and electrode voltages, one looks for the potential distribution V(r,z):

 $\partial^2 V(r,z)/\partial z^2 + (1/r) \partial V(r,z)/\partial r + \partial^2 V(r,z)/\partial r^2 = 0$

(boundary conditions: $V(r,z)=V_{ext}$, the applied voltages).

As $n(r,z) = \sqrt{V(r,z)}$ is the electrostatic index of refraction, the ray equation giving the trajectory r=r(z) can then be integrated:

 $2V(r,z) \frac{\partial^2 r(z)}{\partial z^2} + \frac{\partial V(r,z)}{\partial z} \frac{\partial r(z)}{\partial z} + \frac{(r/2)\partial^2 V(r,z)}{\partial z^2} = 0$

(initial conditions: $r=r_0$ (position) and $dr/dz=tan \phi_0$ (slope) at z=0).

Solution yields the output coordinate r=r(L) of the ray. Repeating ray calculations with different tan ϕ_0 , we get the confusion distribution

$r=r(L,r_0)$

of the electron-optics for entrance at $[z=0, r=r_0]$ in the object plane.

Many steps of optimization are repeated on the full image, with slight changes the electrodes shapes and voltages so as to minimize aberrations all over the image-field.

Common PMT structures



Common PMT structures (cont'd)



Types of PMT responses to be considered

•*Charge* (or integral) response:is the number N of electrons collected at the anode for one photoelectron impinging on the first dynode (*multiplier chain response*) or one photon detected by the photocathode (*PMT response*). Extension: to a light pulse containing R photons (*scintillation counter response*). For all statistics, the random variable N is characterized by mean $\langle N \rangle$ and variance $\sigma_N^2 = \langle [N - \langle N \rangle]^2 \rangle$.

•Impulse (or current) response in the time domain: is the timedependent output SER (single electron response) for an electron at t=0 on the first dynode. We will calculate: the mean $\langle SER(t) \rangle$ and variance $\sigma_{SER}^2(t)$, the output current i(t) for 1 or R photon/s detected at t=0, and the response to an optical signal I(t) carrying R photons. A variant is the *correlation response*.

•Response in the *frequency domain*: is the transfer function $F(\omega)$ of the PMT, given by the mean and noise spectral density $s(\omega)$ for different inputs.

•*Time localization*: is about time measurements, i.e., the delay T between a short light pulse detected at t=0 and the output anode pulse, where $\langle T \rangle$ is the mean delay and the variance σ_T^2 yields the accuracy of time localization supplied by the PMT.



Charge response

multiplication at the dynode (fixed input)

for 1 impinging electronfor K impinging electrons $\langle m \rangle = \langle n \rangle = g$, $\langle m \rangle = K \langle n \rangle = Kg$, $\sigma_n^2 = g$ $\sigma_m^2 = K\sigma_n^2 = Kg$

model of dynode multiplication (random input)



Applying iteratively these equations we get:			
dynode	mean (N	variance σ_N^2	
1°	\mathbf{g}_1	g_1	
2°	$g_{1}g_{2}$	$g_1g_2 + g_1g_2^2$	
3°	$g_1g_2g_3$	$g_1g_2g_3 + g_1g_2g_3^2 + g_1g_2^2g_3^2$	
4°	$g_1g_2g_3g_4$	$g_1g_2g_3g_4 + g_1g_2g_3g_4^2 + g_1g_2g_3^2g_4^2 + g_1g_2^2g_3^2g_4^2$	
••		• • • • • • • • • • • • • • • • • • • •	
n-th	$g_1g_2\cdots g_n$	$\sum_{i=1,n} g_1 g_2 \cdots g_i \ [g_{i+1} \cdots g_n]^2$	
Whence : $\langle N \rangle = g_1 g_2 \cdots g_n$			
$\sigma_{N}^{2} = \langle N \rangle^{2} \sum_{i=1,n} 1 / (g_{1}g_{2}^{\dots}g_{i})$			

Charge response (cont'd)

Letting all dynode gains g equal except the first g_1 we have:

$$\sigma_{N}^{2} = \langle N \rangle^{2} (1/g_{1}) [1+1/g+1/g^{2}+\dots+1/g^{n-1}]$$
$$\approx \langle N \rangle^{2} / [g_{1}(1-1/g)] = \langle N \rangle^{2} \varepsilon_{A}^{2}$$

where

$$\epsilon_{\rm A}^{2} = g / [g_1(g - 1)]$$

is called the *multiplier variance*, because it just gives the relative variance of the charge, $\sigma^{2}/(N)^{2} - c^{2}$

$$\sigma_{\rm N}^2 / \langle {\rm N} \rangle^2 = \epsilon_{\rm A}^2$$

Charge response (cont'd)

For *exactly* R photoelectrons arriving to the 1st dynode,

$$\langle N_R \rangle = g_1 g_2 \cdots g_n R = \langle N \rangle R$$

 $\sigma_{NR}^2 = \sigma_N^2 R = \langle N_R \rangle^2 \epsilon_A^2 R$

and the relative variance is:

$$\sigma_{NR}^{2}/\langle N_{R} \rangle^{2} = \epsilon_{A}^{2}/R$$

i.e., it improves as 1/R.

Last, consider the R photoelectrons be Poissondistributed, as due to F photons detected by the photocathode with quantum efficiency η , so that $R = \eta F$

photoelectron are emitted.

Charge response (cont'd)

If F is Poisson distributed, also $R = \eta F$ is Poisson distributed, whence $\sigma_R^2 = \langle R \rangle$. The statistical compounding rules are schematized below:



By applying the compounding rules, we get:

 $\langle N_{\rm F} \rangle = \langle \eta F \rangle \langle N \rangle$ $\sigma_{\rm NF}^2 = \langle \eta F \rangle \sigma_{\rm N}^2 + \sigma_{\rm R}^2 \langle N \rangle^2 = \langle N_{\rm F} \rangle^2 (1 + \epsilon_{\rm A}^2) / \langle \eta F \rangle$ As the relative variance carried by the packet of photons F is: $\sigma_{\rm F}^2/\langle F \rangle^2 = 1/\langle F \rangle$ the photomultiplier adds a worsening by a factor: $NF^{2} = [\sigma_{NF}^{2} / \langle N_{F} \rangle^{2}] / [\sigma_{F}^{2} / \langle F \rangle^{2}] = (1 + \varepsilon_{A}^{2}) / \eta$ NF² is called the *noise figure* and, is nicely low .! .! [for example, $\varepsilon_{\Delta}^2 = 0.33$ for g=4 and with $\eta = 0.33$ we have $NF^{2}=41$

Current response



The process of multiplication is now time-dependent and interdynode flights are schematized by a function f(t), the pdf of having a time-of-flight t—t+dt when the electron leaves the i-th-dynode at time t=0. The number of secondaries n_T

again electrons follows the Poisson statistics. Each electrons from the i-th dynode covers the interdynode flight independently

Choices for f(t): Gaussian: $f(t) = [\sqrt{(2\pi)\sigma_t}]^{-1} \exp[-(t - t_0)^2/2\sigma_t^2]$ or polynomial-exponential: $f(t) = [\Gamma(m+1)\sigma_t^{m+1}]^{-1} t^m \exp(-t/\sigma_t)$

A random current will be written as

 $\mathbf{I}(\mathbf{t}) = \mathbf{e} \mathbf{N} \mathbf{f}(\mathbf{t})$

where e is the electron charge, N is the random total charge, f(t) the time-dependence

(for simplicity, we will omit e from now on)

mean current at the (i+1)-th dynode:

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\langle n(t) \rangle = \langle n_T \rangle f(t)
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where n_T is the (random) total number of secondaries $[n_T = \int_{0-\infty} \langle n(t) \rangle dt]$, with mean and variance:

$$\langle n_{\rm T} \rangle = g, \quad \sigma_{\rm nT}^2 = g,$$

and as $f(t) = f_i(t)$ we have:

$$\langle n(t) \rangle = g f_i(t)$$

To compute the variance, we shall introduce the generatlized concept of covariance

Covariance function $K_n(t,t')$: $K_n(t,t') = \langle [n(t)-\langle n(t) \rangle][n(t')-\langle n(t') \rangle] \rangle$

for t=t' the covariance becomes the variance, $K_n(t,t) = \sigma_i^2(t)$; and, if K is dependent on t-t'= τ only, it becomes the correlation function $\rho(\tau)$.

For a process with a sequence of n_T independent electrons with arrival times f(t), one can find that the covariance is:

$$\mathbf{K}_{\mathrm{n}}(t,t') = \left[\left(\sigma_{\mathrm{n}T}^{2}/\langle \mathbf{n}_{\mathrm{T}}\rangle\right) + \langle \mathbf{n}_{\mathrm{T}}\rangle - 1\right] \mathbf{f}(t) \ \delta(t'-t)$$

where $\delta(t'-t)$ is the Dirac delta-function.

For a Poisson distributed n_T , we have:

 $K_n(t,t') = g f(t) \delta(t'-t)$





This response is the SER (*single-electron-response*). Mean value is:

 $\langle SER(t) \rangle = g_1 g_2 \cdots g_n f_1(t) * f_2(t) * \cdots * f_n(t)$ Also, as $\langle N \rangle = g_1 g_2 \cdots g_n$: $\langle SER(t) \rangle = \langle N \rangle \quad f_1(t) * f_2(t) * \cdots * f_n(t)$

The covariance is:

$$K_{SER}(t,t') = \sum_{i=1,n} g_1 g_2 \cdots g_i [g_{i+1} \cdots g_n]^2 f_1(t) * f_2(t) * \cdots * f_i(t) * [f_{i+1}(t) * \cdots * f_n(t)] [f_{i+1}(t') * \cdots * f_n(t')]$$

and the SER variance (for t=t') is :

$$\sigma_{\text{SER}}^{2}(t) = \sum_{i=1,n} g_{1} \cdots g_{i} [g_{i+1} \cdots g_{n}]^{2} f_{1}(t) * f_{2}(t) * \cdots * f_{i}(t) * [f_{i+1}(t) * \cdots * f_{n}(t)]^{2}$$

Rigid SER approximation: $f_{1}(t)_{*} \cdots_{*} f_{i}(t)_{*} [f_{i+1}(t)_{*} \cdots_{*} f_{n}(t)]^{2} \approx [f_{1}(t)_{*} f_{2}(t)_{*} \cdots_{*} f_{n}(t)]^{2}$ then we have: $K_{SER}(t,t') \approx \sum 1/(g_{1}g_{2} \cdots g_{i}) \langle SER(t) \rangle \langle SER(t') \rangle$ $= \epsilon_{A}^{2} \langle SER(t) \rangle \langle SER(t') \rangle$ $\sigma_{SER}^{2}(t) \approx \epsilon_{A}^{2} \langle SER(t) \rangle^{2}$



An example of the accuracy of the rigid SER approximation for n=12 stages and a Gaussian time-of-fight distribution



Noise equivalent circuit of the dynode multiplying chain (top) and of the photomultiplier (bottom). Noisegenerators N_i sum to the mean signal a white noise of intensity equal to the mean signal

Case of a fixed number R of photoelectrons in a single short pulse with $\langle L(t) \rangle = \langle F \rangle \Phi(t)$:

 $\langle I(t) \rangle = R f_0(t) * \langle SER(t) \rangle$

 $K_{I}(t,t') = R f_{0}(t)\delta(t'-t)_{*}*K_{SER}(t,t') \approx R f_{0}(t)*\langle SER(t) \rangle \langle SER(t') \rangle \epsilon_{A}^{2}$

 $\sigma_{\rm I}^2(t) \approx {\rm R} \ {\rm f}_0(t) * \langle \ {\rm SER}(t) \rangle^2 \epsilon_{\rm A}^2$

Also in this case, the relative variance of the response, given by:

 $\sigma_{I}^{2}(t)/\langle I(t) \rangle^{2} = [f_{0}(t) * \langle SER(t) \rangle^{2} \varepsilon_{A}^{2}] / R [f_{0}(t) * \langle SER(t) \rangle]^{2}$

can be approximated, on the rigid SER approx., to:

 $\sigma_{\rm I}^2(t) / \langle I(t) \rangle^2 \approx \epsilon_{\rm A}^2 / R$



Some samples of the anode response of the photomultiplier for R=10, 20 (top), 50, 100 (bottom) photoelectrons emitted at t=0, calculated by a Montecarlo method. Horizontal scale: time in ns; vertical scale: current in relative units

Correlation response

The correlation function $\rho(\tau)$ of a light field L(t) is defined as:

$$\rho_{L}(\tau) = \int_{t = 0-\infty} L(t) L(t+\tau) dt$$

If L(t) is a train of photons occurring at random times t_k , i.e.:

$$L(t) = h\nu \sum_{k=-\infty, +\infty} \delta(t-t_k)$$

developing ρ_L the correlation is found as:

$$\rho_{\rm L}(\tau) = (h\nu)^2 \{ F \,\delta(\tau) + F^2[1 + \gamma(\tau)] \}$$

where $\gamma(\tau)$ is the reduced correlation function,

 $\gamma(\tau) = \left\langle \left[L(t) - \left\langle L(t) \right\rangle \right] \left[L(t + \tau) - \left\langle L(t + \tau) \right\rangle \right] \right\rangle / \left\langle L(t) \right\rangle^2$

Correlation response (cont'd)

The output current correlation $\rho_I(\tau)$ of I(t), can then be computed and the result is :

$$\rho_{I}(\tau) = (\eta e)^{2} F \rho_{\Delta SPR}(\tau) + (\eta e)^{2} F^{2} [1 + \gamma(\tau)] * \rho_{\langle SPR \rangle}(\tau)$$

where

$$\rho_{\langle SPR \rangle}(\tau) = \int_{t=0-\infty} \langle SPR(t) \rangle \langle SPR(t+\tau) \rangle dt$$

is the autocorrelation associated with the mean waveform $\langle SPR(t) \rangle = f_0(t)^* \langle SER(t) \rangle$ of the single photon response, and

$$\rho_{\Delta SPR}(\tau) = \rho_{\langle SPR \rangle}(\tau) + \int_{t=0-\infty} K_{SPR}(t,t+\tau) dt$$

is the autocorrelation of the SPR fluctuations, contained in the covariance K given by :

$$K_{SPR}(t,t+\tau) = \varepsilon_A^2 f_0(t) * \langle SER(t) \rangle \langle SER(t') \rangle$$

Correlation response (cont'd)



Correlation response (cont'd)

When the correlation times are much longer, or $f_0(t)$ is short compared to the SER, a rigid SPR approximation is adequate and the previous equations become:

 $\rho_{\Delta SPR}(\tau) = (1 + \epsilon_A^{2}) \rho_{\langle SPR \rangle}(\tau)$

 $\rho_{I}(\tau) = (\eta e)^{2} \left(1 + \epsilon_{A}^{2}\right) F \rho_{\langle SPR \rangle}(\tau) + (\eta e)^{2} F^{2} \left[1 + \gamma(\tau)\right]_{*} \rho_{\langle SPR \rangle}(\tau)$

By deconvolution of $\rho_I(\tau)$ with $\rho_{\langle SPR\rangle}$, light correlations $\gamma(\tau)$ waveforms down to about 10% of SER duration can be recovered

Frequency Domain response

The transfer function $F(\omega)$ is the Fourier transform of the δ -response $U_{\delta}(t)$, and the noise power spectrum $S^2(\omega) = di^2/d\omega$ is the Fourier transform of $K_U(\tau)$:

frequency response:
$$F(\omega) = F\{\langle SER(t) \rangle \},\$$

for a Gaussian waveform : $F(\omega) = \langle N \rangle \exp(-\omega^2 n \sigma_t^2/2 + i \omega n t_0)$, whence a high frequency cutoff (at -3dB) of $f_T = 0.83/2\pi\sigma_t \sqrt{n} = 0.13/\sigma_t \sqrt{n}$.

noise power spectral density: $S^{2}(\omega) = 2e (I_{ph}+I_{d}) (1+\epsilon_{A}^{2}) |F(\omega)|^{2}$

In the passband, the PMT is close to the quantum limit for $I_{ph0} \ge I_d$, i.e. for an input power larger than $P_{i0} \ge (h\nu/\eta e)I_d$. In this case, the S/N is:

$$(S/N)^2 = I_{ph} / 2eB(1+\epsilon_A^2) = P/2(h\nu/\eta)B(1+\epsilon_A^2)$$

noise figure: NF²= $(1+\varepsilon_A^2)/\eta$.

 P_i is very small, even for small η , because of the very minute dark currents I_d of the photocathode, for ex., for $I_d = 1$ f A and $\eta = 0.1$, quantum-limited performance is at $P \ge P_{i0} = 20$ fW and is over the full frequency band.

Time Sorting and Measurements



Time sorting can be simply performed by a threshold crossing at a suitable level S_o and taking T_o as the timing information carried by the signal

Time Sorting and Measurements (cont'd)

Time variance for crossing of a fixed amplitude threshold:

$$\sigma_{\rm T}^2({\rm T_o}) = \sigma_{\rm S}^2({\rm T_o})/\langle | dS/dt |_{t={\rm To}} \rangle^2$$

for the time measurement on SER:

$$\sigma_{T1}^{2} = \epsilon_{A}^{2} \langle SER(t) \rangle^{2} / [d/dt \langle SER(t) \rangle]^{2}$$

and for the time measurement on a fast pulse:

For the time measurement on SER:

 $\sigma_{T\delta}^{2} = [(1 + \varepsilon_{A}^{2})/R] \{ f_{o}(t) * \langle SER(t) \rangle \}^{2} / \{ d/dt [f_{o}(t) * \langle SER(t) \rangle] \}^{2}$

Example: for Gaussian-distributed time-of-flights $f_0, f_1, ..., f_n$, each with a variance σ_t , using a threshold at the maximum slope of the SER we have:

$$\sigma_{T\delta}^2 = [(1 + \epsilon_A^2)/R](n+1)\sigma_t^2$$

Time Sorting and Measurements (cont'd)

The variance of the SER centroid, as the weighted sum of interdynode-flight uncertainties is:

$$\sigma_{\mathrm{Tc}}^{2} = \sigma_{\mathrm{t0}}^{2} + (1/g_{1})\sigma_{\mathrm{t1}}^{2} + (1/g_{1}g_{2})\sigma_{\mathrm{t2}}^{2} + \dots + (1/g_{1}g_{2}\dots g_{\mathrm{n}})\sigma_{\mathrm{tn}}^{2}$$

and for equal $\sigma_{ti} = \sigma_t$:

$$\sigma_{\rm T}^2 \approx \sigma_{\rm t0}^2 + \varepsilon_{\rm A}^2 \, \sigma_{\rm t}^2$$

and in the case of $R = \langle \eta F \rangle$ photoelectrons:

$$\sigma_{Tc}^{2} = [\sigma_{t0}^{2} + (1/g_{1})\sigma_{t1}^{2} + (1/g_{1}g_{2})\sigma_{t2}^{2} + \dots + (1/g_{1}g_{2}\dots g_{n})\sigma_{tn}^{2}]/R$$
$$= [\sigma_{t0}^{2} + \varepsilon_{A}^{2}\sigma_{t}^{2}]/R$$

Written as $\sigma_{Tc}^2 = [1 + \epsilon_A^2] \sigma_t^2 / R$ the centroid variance is smaller by a factor (n+1) as compared to the fixed-threshold crossing.